

Chequerboard Phased Array Feed Testing for ASKAP

**A. Chippendale, J. O'Sullivan, J. Reynolds, R. Gough, D. Hayman,
S. Hay, R. Shaw, R. Qiao**

BYU Phased Array Workshop

3 May 2010

Photo by David McClenaghan



Analog Systems Team

- Russell Gough (IPT leader)
- John O'Sullivan (IPT scientist)
- Steve Barker (IPT PM)
- **Electromagnetic design of FPA, calibration system, and testing**
 - Stuart Hay
 - Rong-Yu Qiao
 - Francis Cooray
 - Doug Hayman
 - Aaron Chippendale
- **Mechanical design**
 - Deszo Kiraly
 - Russ Bolton
 - Paul Doherty
 - Eliane Hakvoort
 - Workshop staff (3 EFT)
- **Low-noise amplifier design**
 - Rob Shaw
- **Receiver electronics**
 - K. Jeganathan
 - Simon Mackay
 - Suzy Jackson
 - Yoon Chung
 - Peter Axtens (PSUs)
- **Plus consultation services of:**
 - Pat Sykes
 - Michael Brothers
 - Matt Shields
 - Mark Bowen
 - Santy Castillo
 - Henry Kanoniuk
 - Les Reilly

Key Topics

- **Overview** of chequerboard development
- **Results** from the 5×4 prototype
- **Progress** towards ASKAP and SKA

Overview – Why Phased Array Feeds?

- **Key SKA requirement**

- sensitive observations of the sky over large areas (surveys)
- collect more information from the sky per antenna

- **Achieve target SKA survey speed with**

- order of magnitude fewer antennas
- 20% to 50% lower cost
Chippendale et al. 2007, SKA Memo 92

- **Optimise performance for science**

- control beam/FOV shape
- control instrumental polarisation
- control spillover
- control baseline ripple
- mitigate interference

Overview – ASKAP Receiver Specifications

Focus package

- frequency range..... 700 MHz to 1800 MHz
- field of view.....30 deg²
- T_{sys} < 50 K (aspire to 35 K)
- efficiency η unspecified (but ~65%)
- PAF Ports.....188 (94×2 pol.)
- weight.....< 200 kg
- continuum dynamic range..... 10^6
- spectral dynamic range..... 10^6
- post calibration polarisation purity..... -30 dB

Analog system cost estimate per antenna

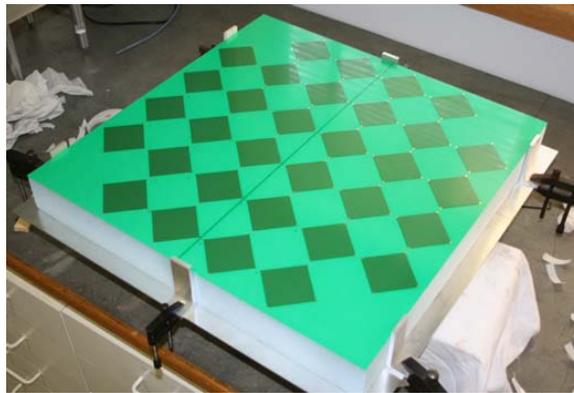
- focus (gain + band select).....\$150 k
- pedestal (downconversion).....\$120 k
- cooling..... \$15 k

Overview – Connected Element Array Evolution



5 port

Hay et al. 2007, EuCAP

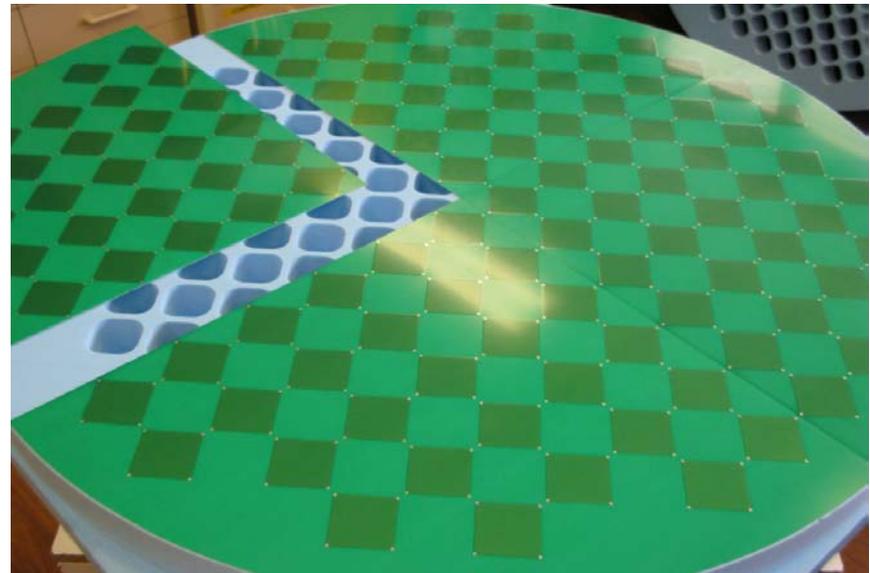


40 port

approximately
self complementary

$$Z \approx 377 \Omega$$

188 port

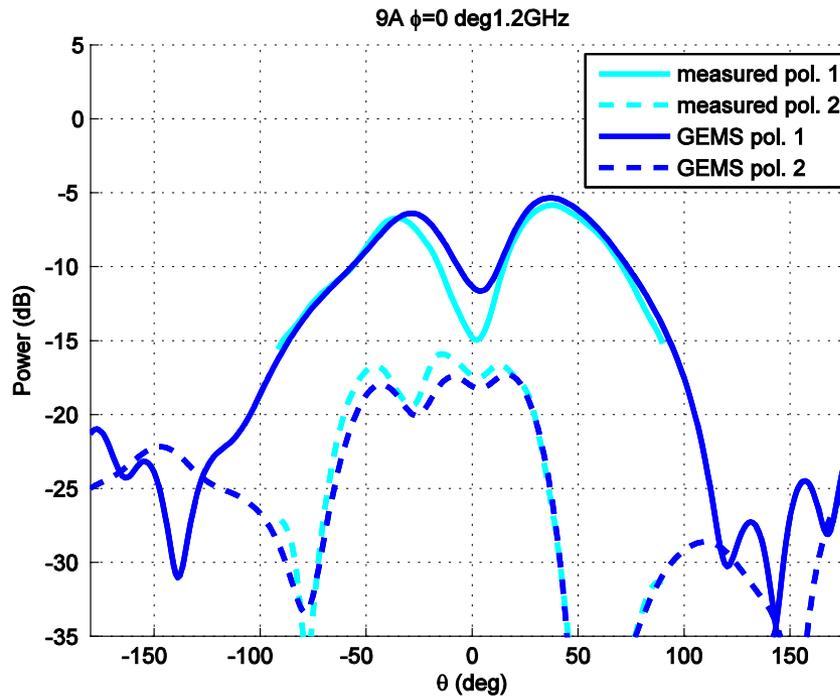


Advantages

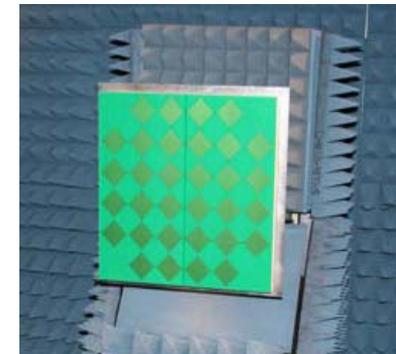
- lower ohmic loss
- lower polarisation variation
- easy manufacture
- cryo-cooling integration potential

Testing – Power Patterns

Simulated and measured port power patterns agree



40 port

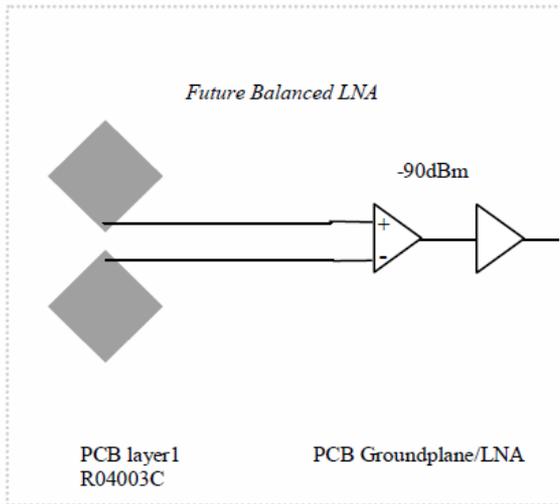


Measurement Setup

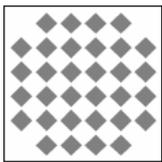
Hay et al. 2007, EuCAP

Testing – 5x4 Rx block diagram

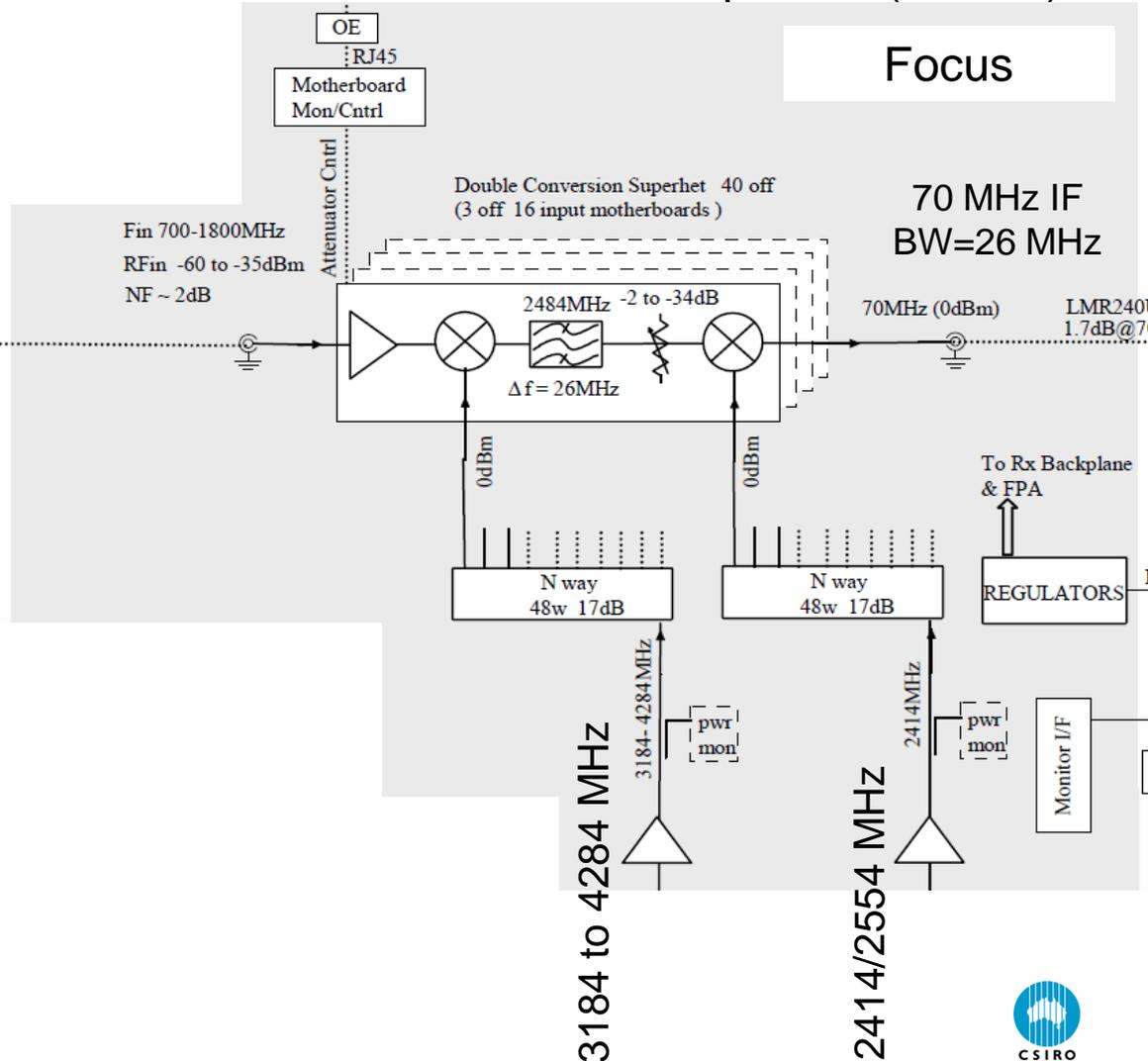
SCA Element (1 of 40)



40 Element SCA
The "5x4" Planar Array



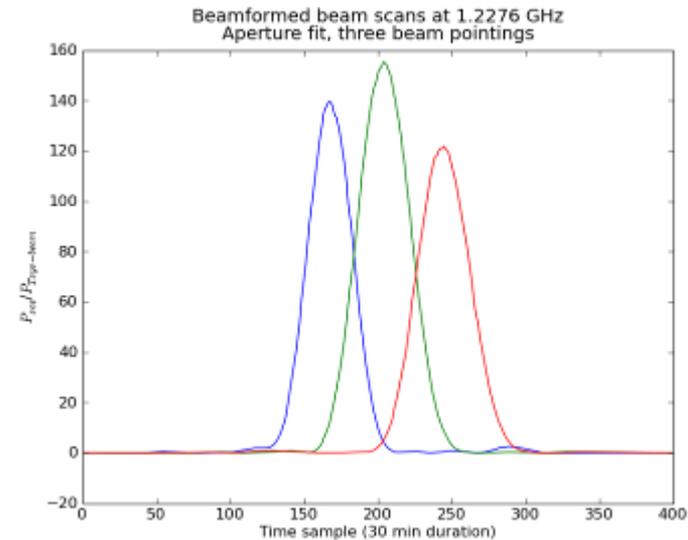
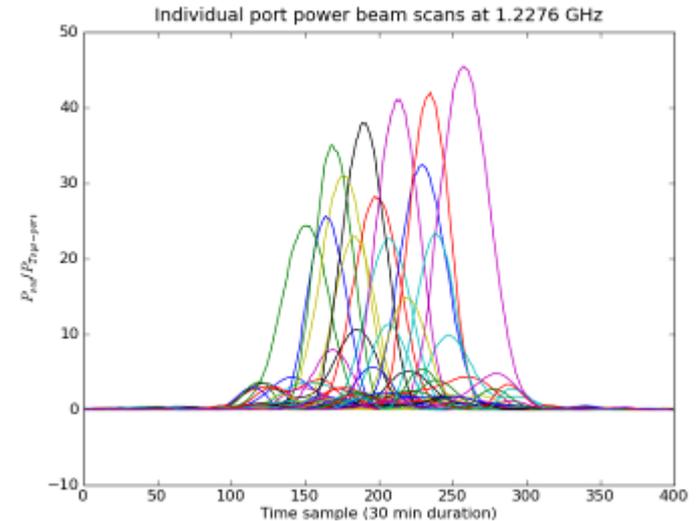
Double Conversion Superhet. (1 of 48)



Testing – Parkes Single Dish

Single dish

- 0.875 MHz BW
- samples \rightarrow disk
- software correlation
- $T_{sys}/\eta \sim 175 \text{ K +/- ?}$
- driftscans:
 - beamform on GPS (L2 1227.6 MHz)
 - measure T_{sys}/η on Virgo A
- beamformed:
 - $\eta \sim 4 \times$ element η
 - $T_{sys} \sim 0.8 \times$ element T_{sys}



Testing – Parkes Interferometer



Testing – Measuring T_{sys}/η



beamformed
PAF voltage

$$v_{beam} = \sum_i w_i v_i$$



64 m reference
voltage

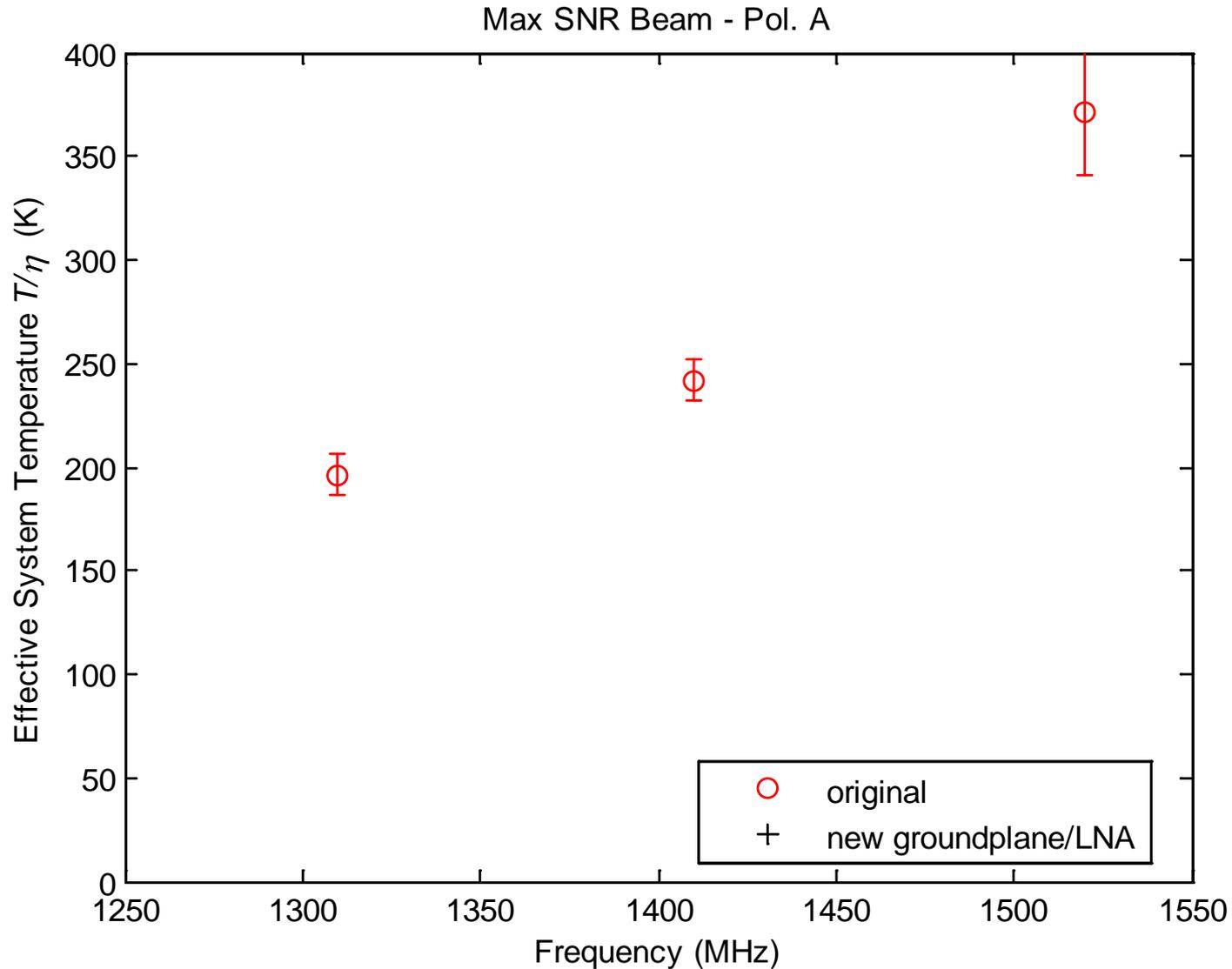
r

auto + cross-correlations on/off point source
 → high SNR measure of T_{sys}/η requiring few assumptions

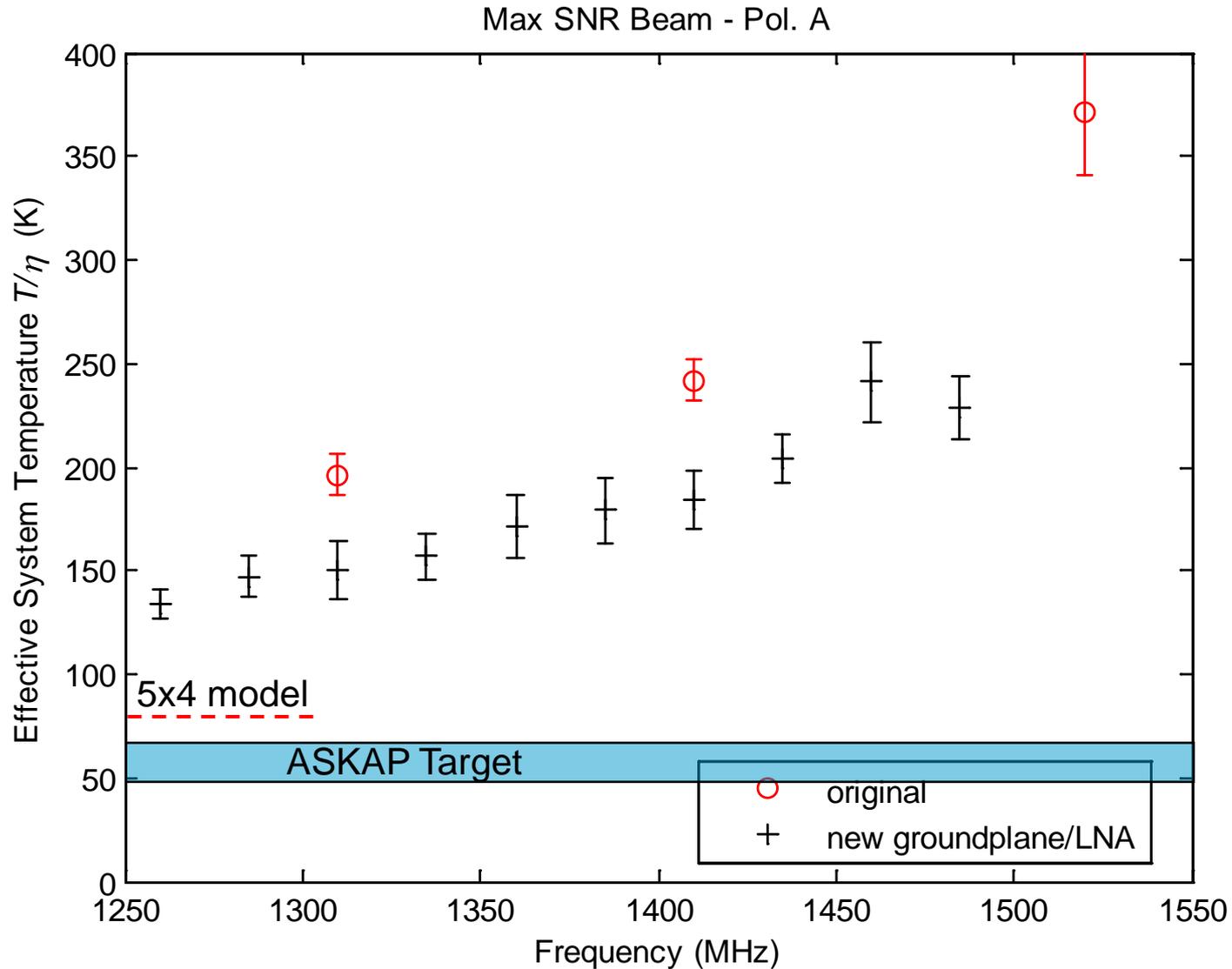
maximise gain $[w_i] = [v_i^*]$

maximise SNR $\mathbf{w} = \mathbf{R}^{-1} \mathbf{v}^*$ covariance matrix $[R_{ij}] = [\langle v_i v_j^* \rangle]$

Testing – Measuring T_{sys}/η



Testing – Modified Groundplane and New LNA



Testing – Measuring Aperture Plane Array T_{sys}



Beamformed on hot microwave absorber

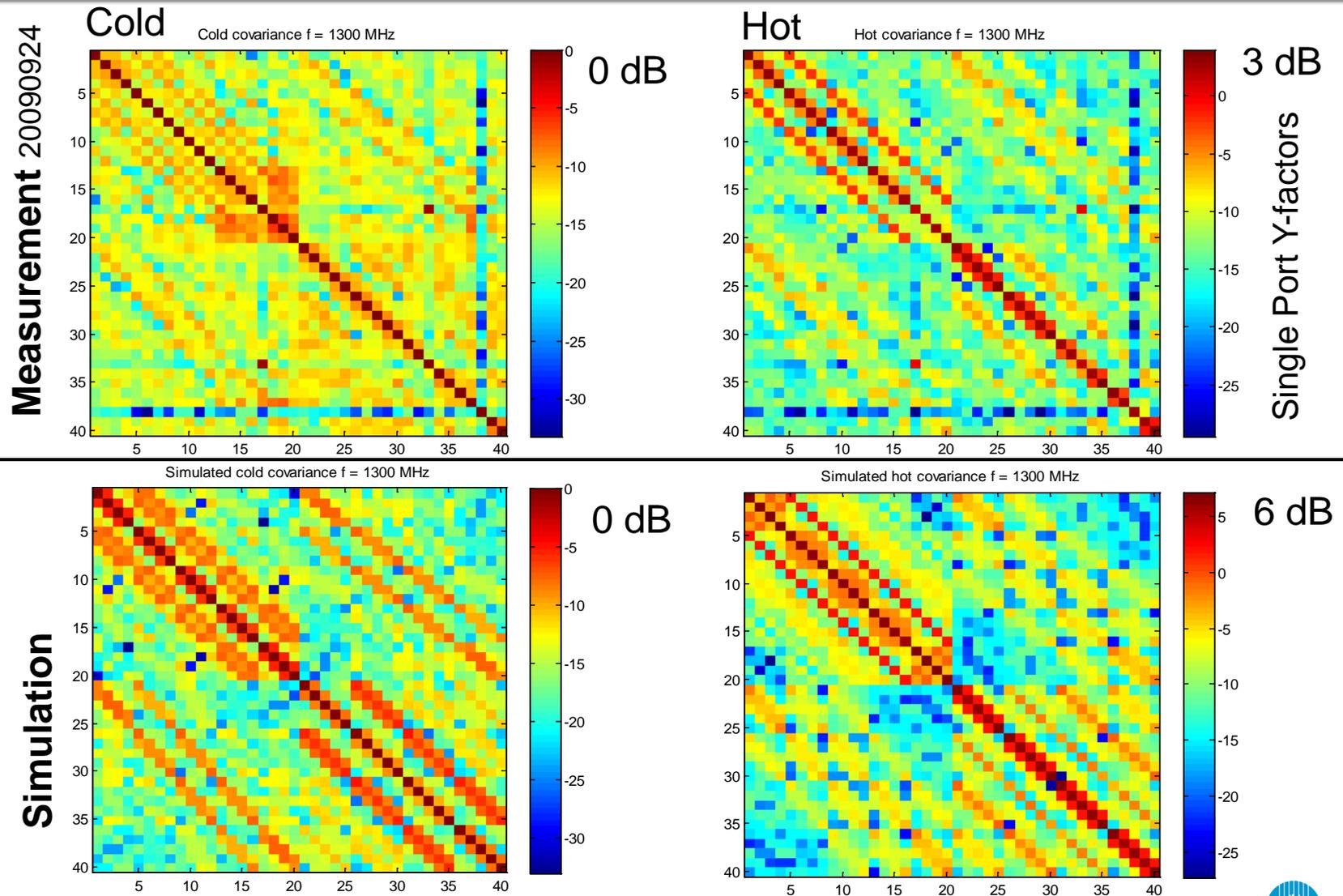
Measured Y-factor from cold sky to hot load

T_{sys} better than PAF operation if all ports used in beamforming

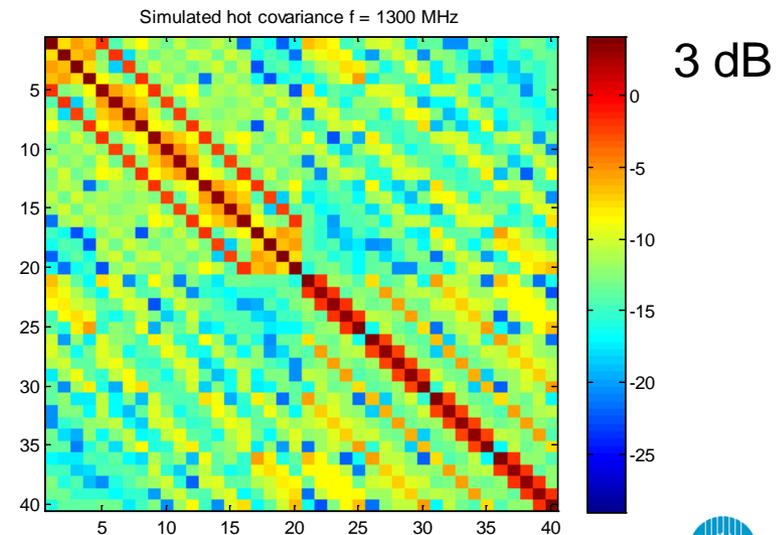
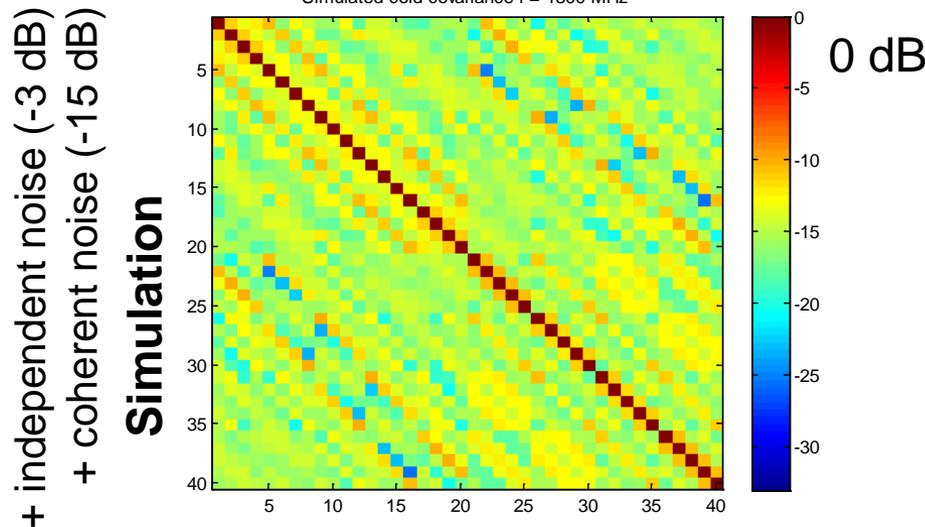
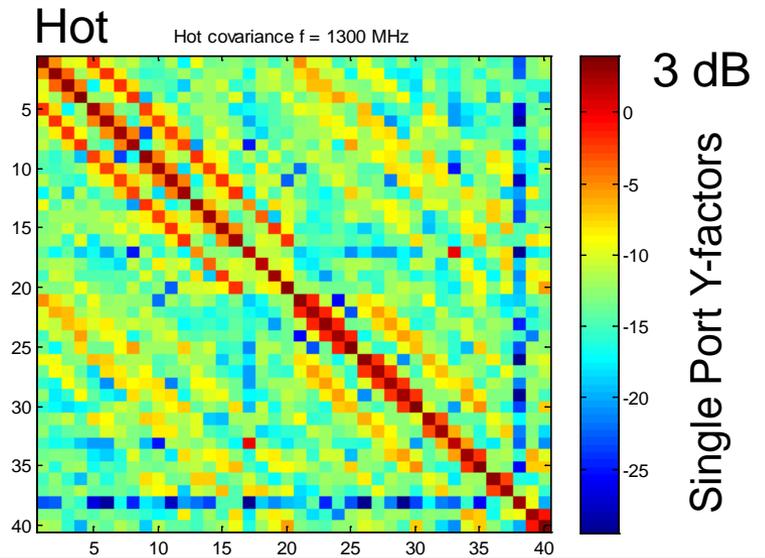
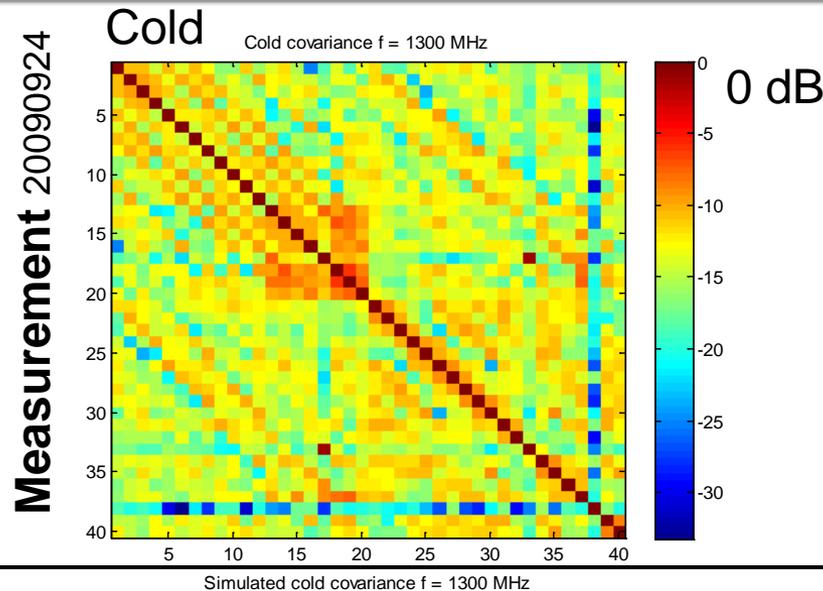
Poor T_{sys} if central 4 ports used

Finite absorber size was accounted for

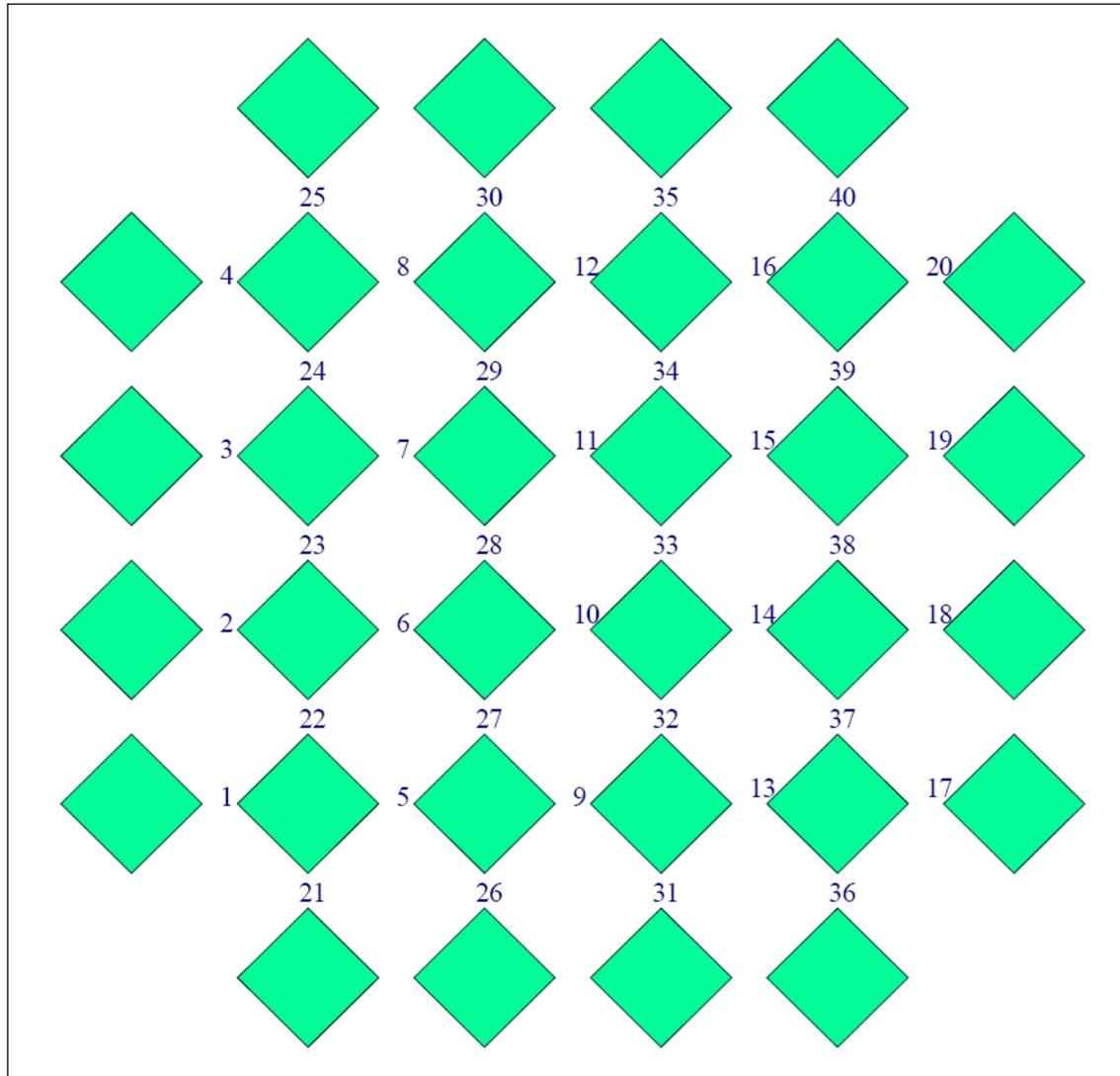
Testing – Measured vs. Simulated Covariance Normalised to Cold Covariance



Testing – Measured vs. Simulated Covariance Normalised to Cold Covariance



Testing – 5×4 Port Numbering



5X4 Array

Front View in Simulation
Rear View in Measurement

Elements 1- 20 polarity "H"

Elements 21 - 40 polarity "V"

Testing – Hunting Unmodelled Noise

First order candidates

1. Mismatch of active array impedance
 - LNA models inaccurate?
 - ASTRON measurements to improve models
2. Self interference
 - Noise observed at IF from switch-mode power supplies
 - Now using linear supplies for 5×4 prototype
 - Digital noise from back-end?

Second order sources discovered

1. LO distribution noise
 - Not significant if RF input level to mixer set appropriately
2. Blockage and scattering
 - Not expected to be significant if LNA models correct
 - Ongoing modelling by Stuart Hay

Testing – Practical Issues

Practical sources of angst

- press-fit connectors + planar couplers in LO distribution
- insufficient headroom in LO distribution
- unreliable stepped attenuator control
- broken tracks in regulator PCB
- aggressive timeline
- poor heat dissipation from focus package

Lessons learned

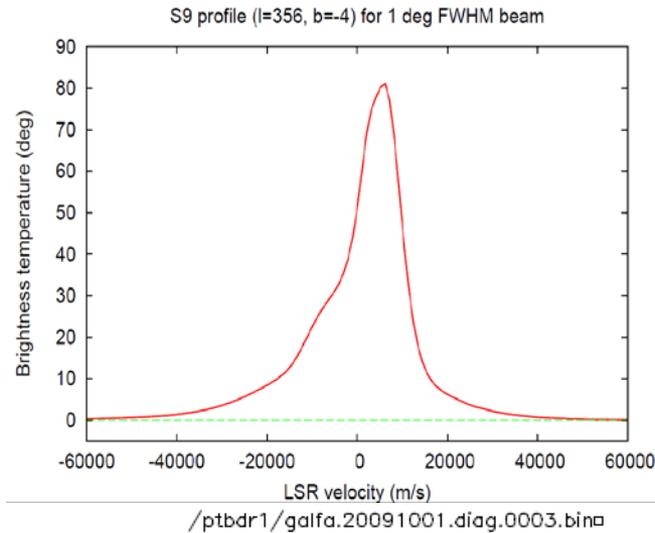
- keep prototype as simple as possible
- make minimum sufficient demonstration of key technology only
- allow more time to cogitate results between tests
- plan testing during design
- test subsystems before integration
- integrate early and often
- don't unlearn lessons learned (see Doug Hayman's presentation)

Future directions (parallel)

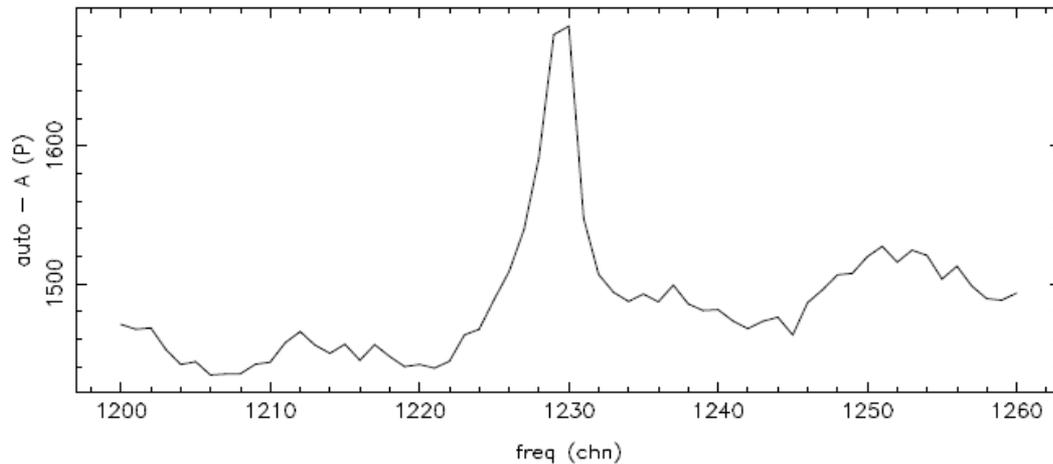
- refurbish and retest existing 5x4 array + receiver
- implement simplified 4-port receiver
 - 5x4 array + BPF + gain + 4 port oscilloscope

Testing – HI Spectrum from Single Port

S9



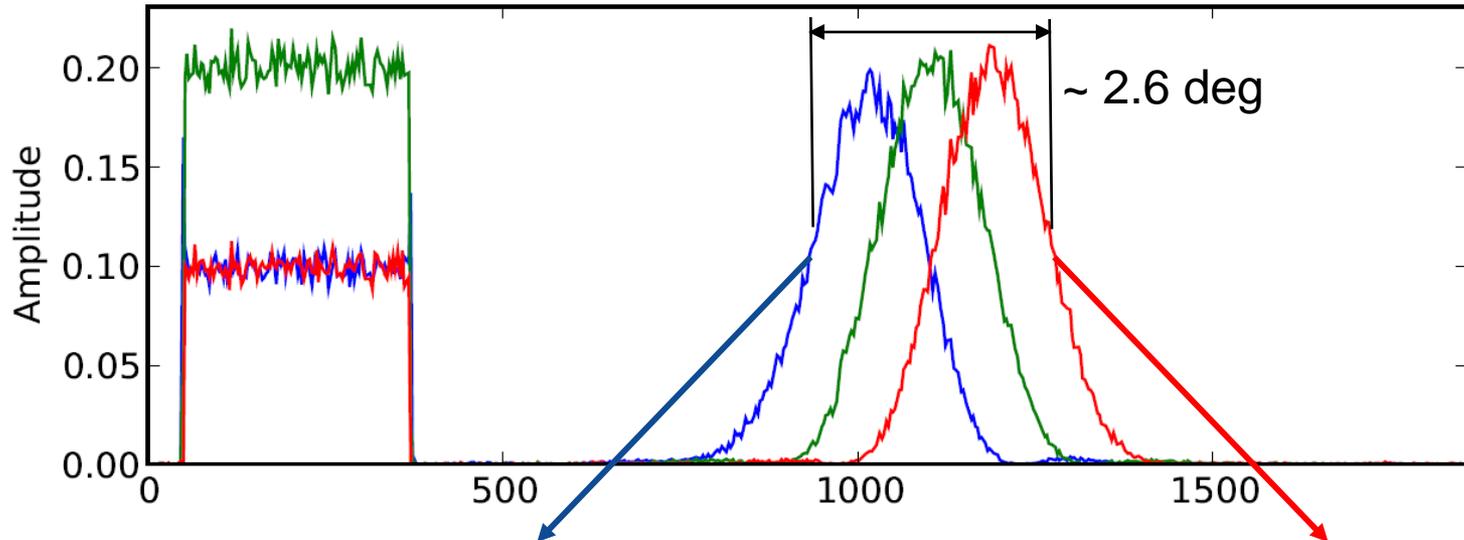
LAB Galactic HI Survey
Smoothed to 1 deg
P. Kalberla et al., 2005



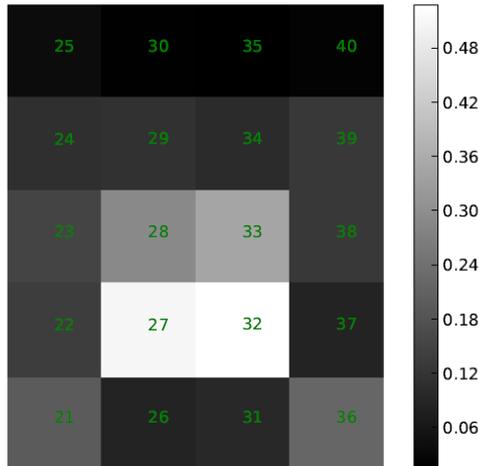
5×4 PAF
Single Port
Spectrum

Testing – Beamformed Power Pattern

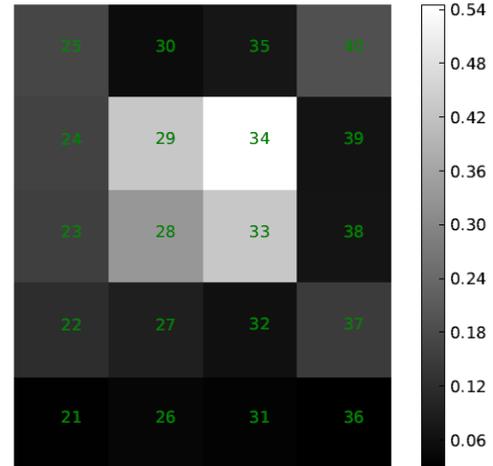
1226+023 epoch: 2008/11/4 23:10:07 UTC CFM beams pol B



1226+023 epoch: 2008/11/4 23:10:07 UTC
CFM beam B pre V-pol. weights



1226+023 epoch: 2008/11/4 23:10:07 UTC
CFM beam B post V-pol. weights



Testing – Next Steps

June 2010

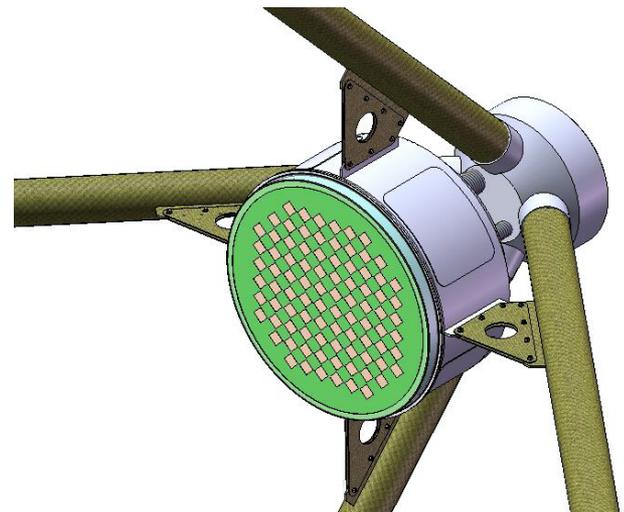
- Hot/Cold measurements at Parkes with linear power supplies

August 2010

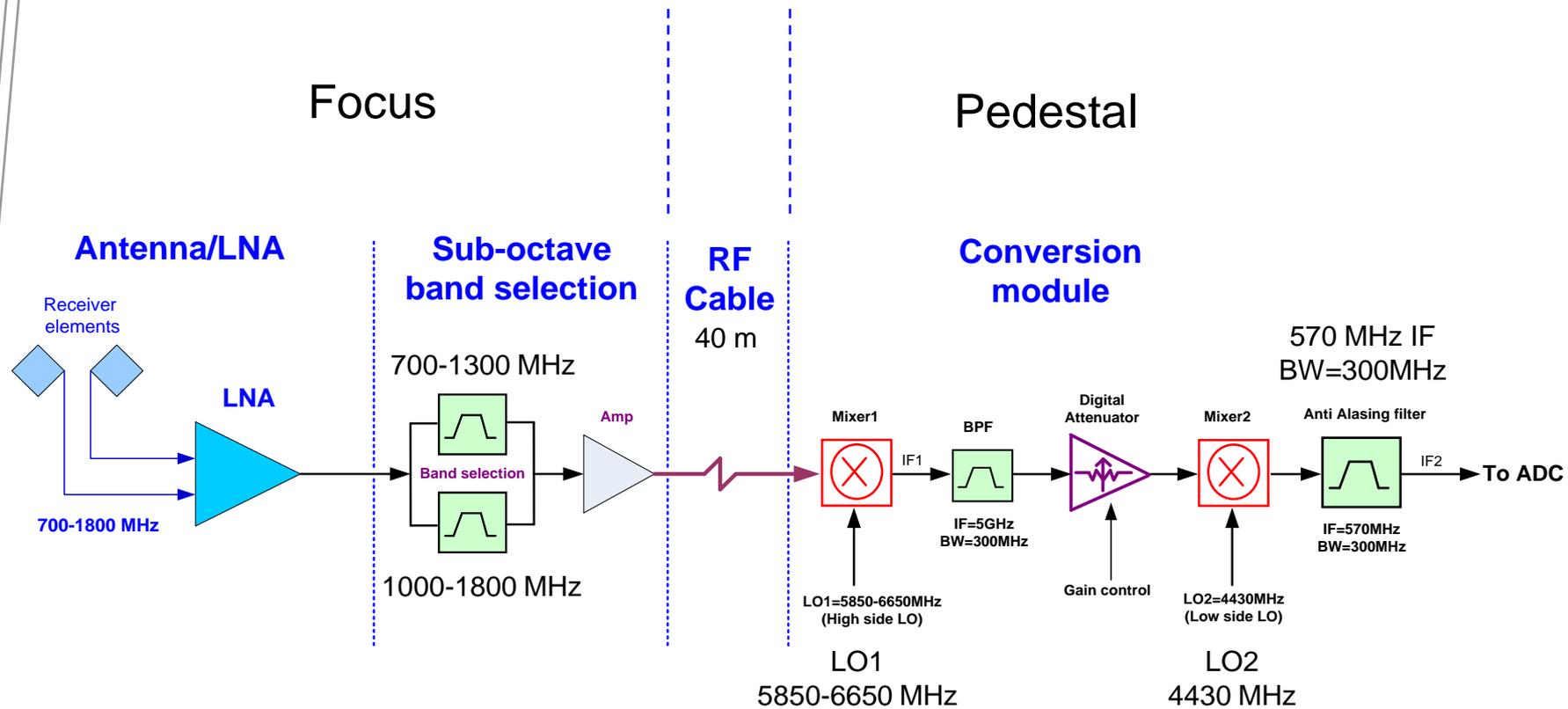
- Repeat hot/cold measurements with simplified back-end

Late 2010

- Start testing 188 port ASKAP PAF

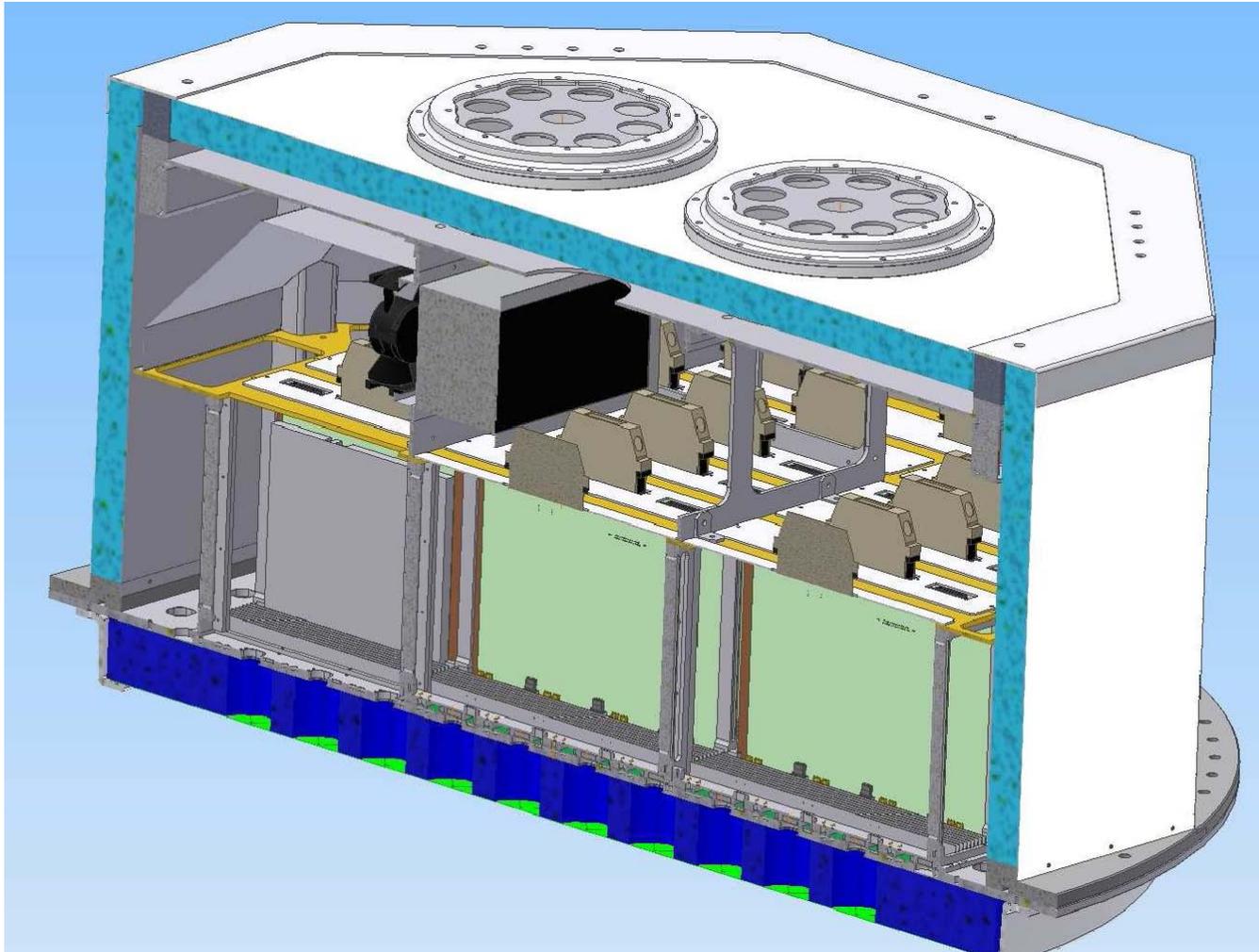


Progress – ASKAP Receiver Block Diagram



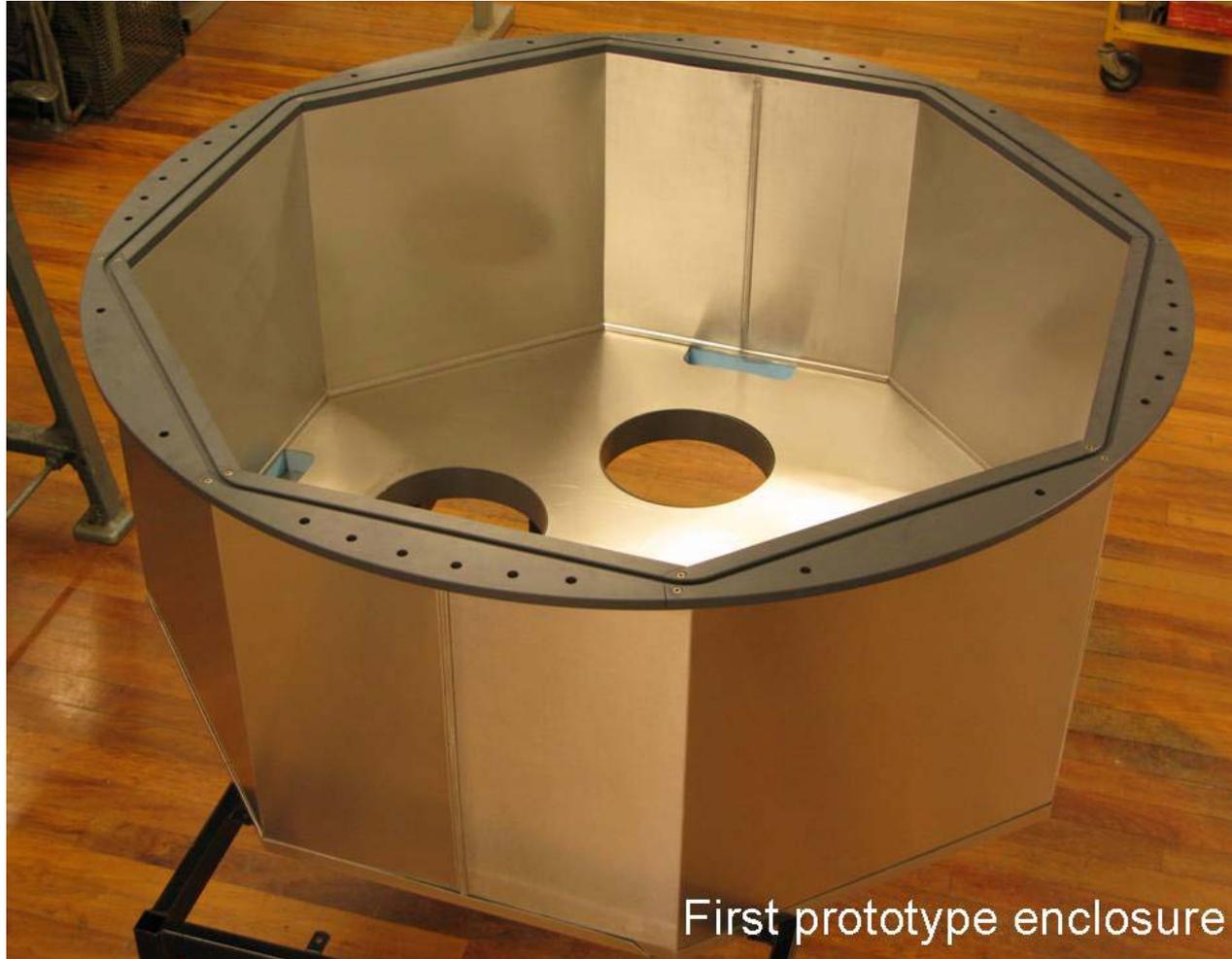
Progress – ASKAP Focus Package Assembly

- Designed for rigidity, EMC/weather-tightness, thermal control



Progress – Focus Package Enclosure

- Double Aluminium skin 40mm EPS foam and PVC flange

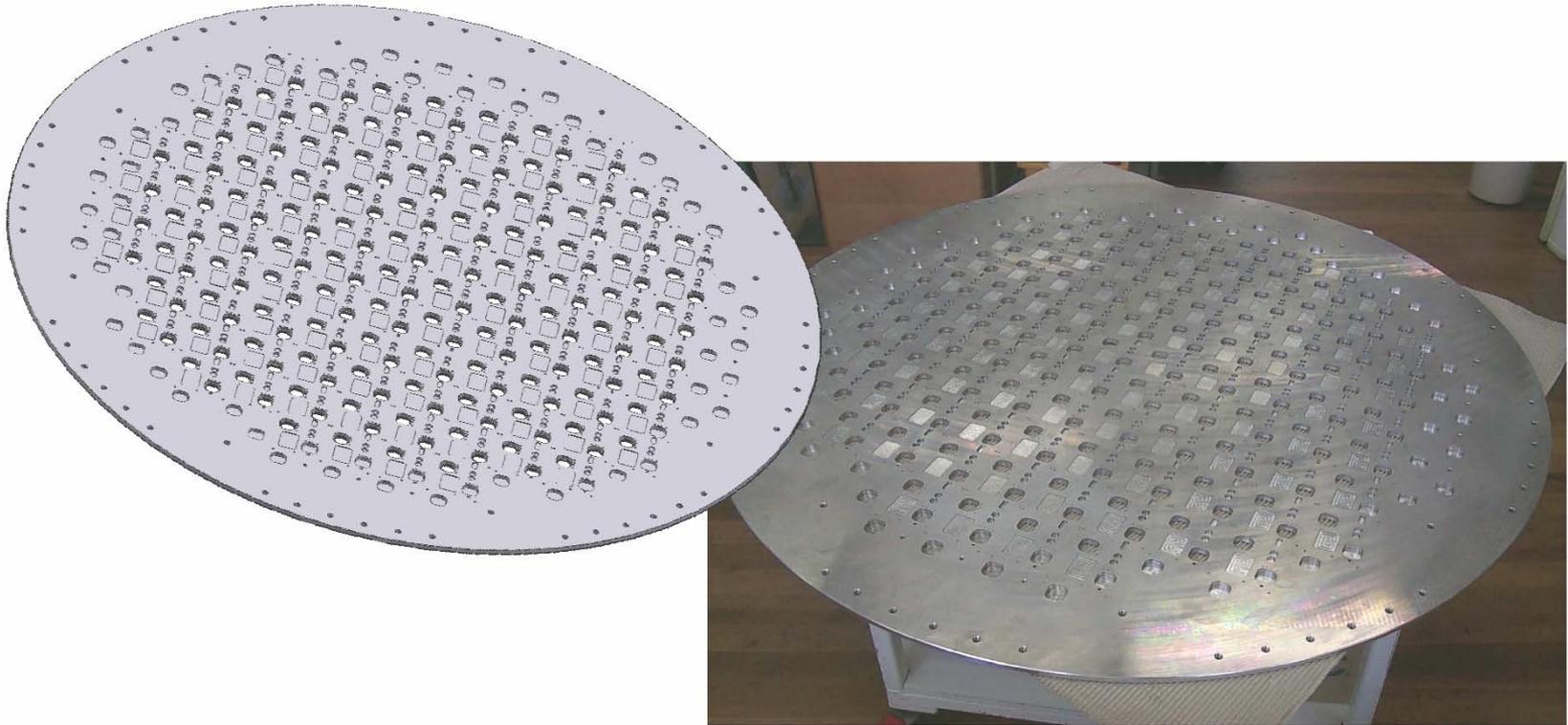


assembly images
c/- Russ Bolton

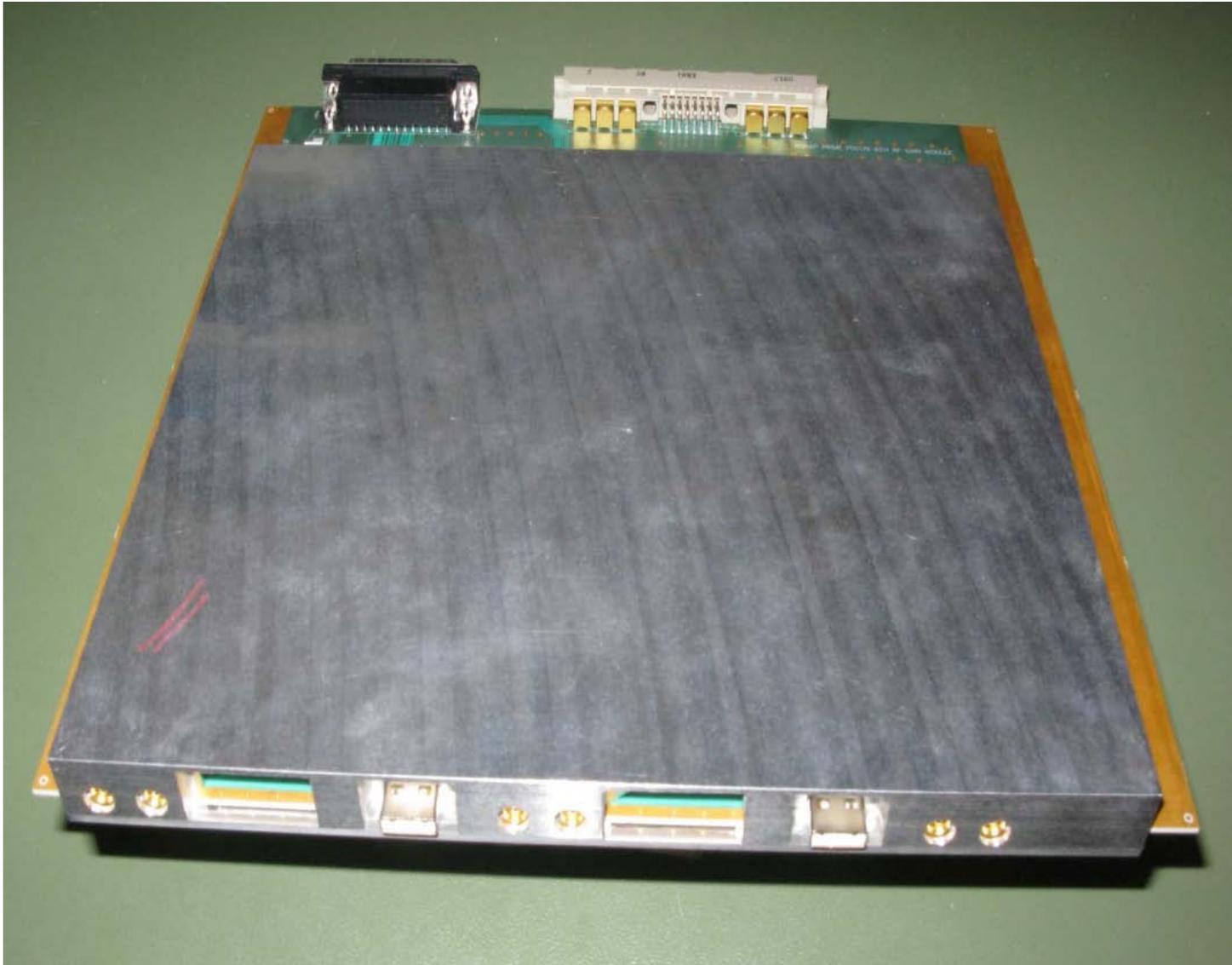
First prototype enclosure

Progress – ASKAP Baseplate

- **Single largest component:** 33 kg weight, 1350 mm diameter
- **Outsourced:** CNC machined 9.5 mm aluminium tooling plate



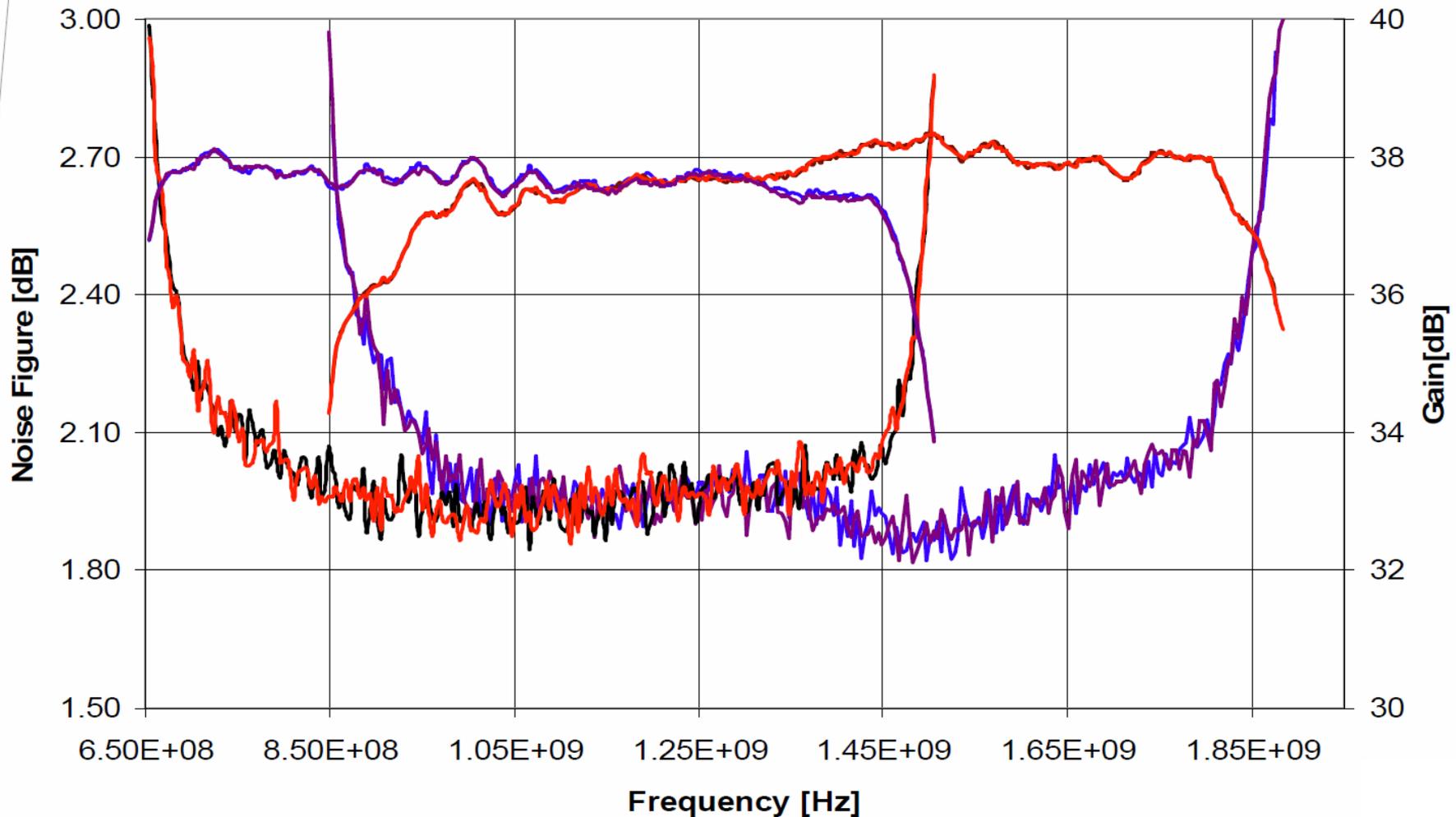
Progress – Gain Card + Shield



Progress – Gain Card EMC Test Linear vs. Switched Power Supply

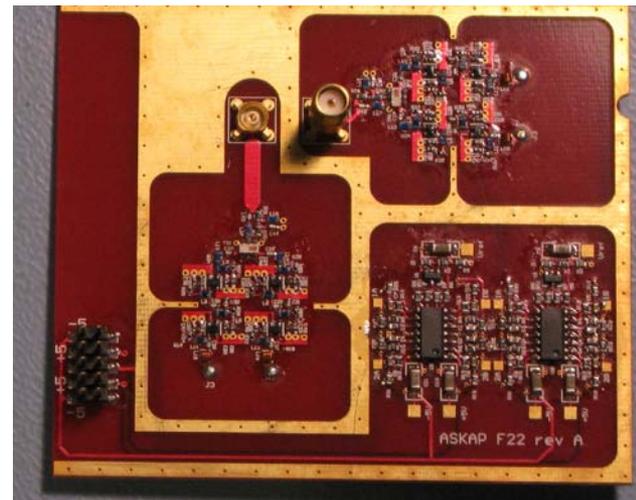
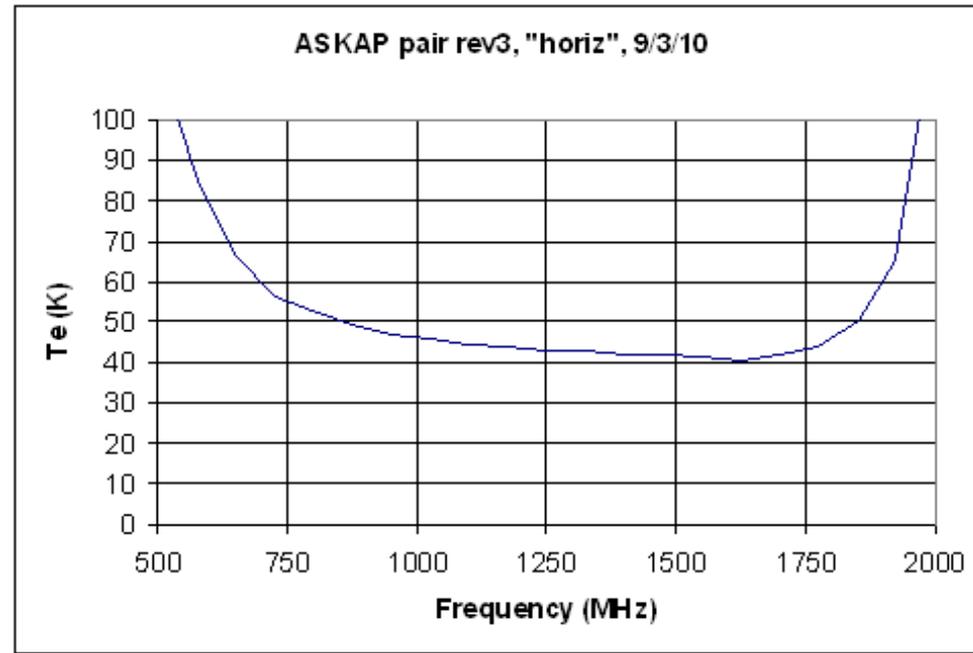
c/- Peter Axtens

Gain-Noise [Amp-Filter module]



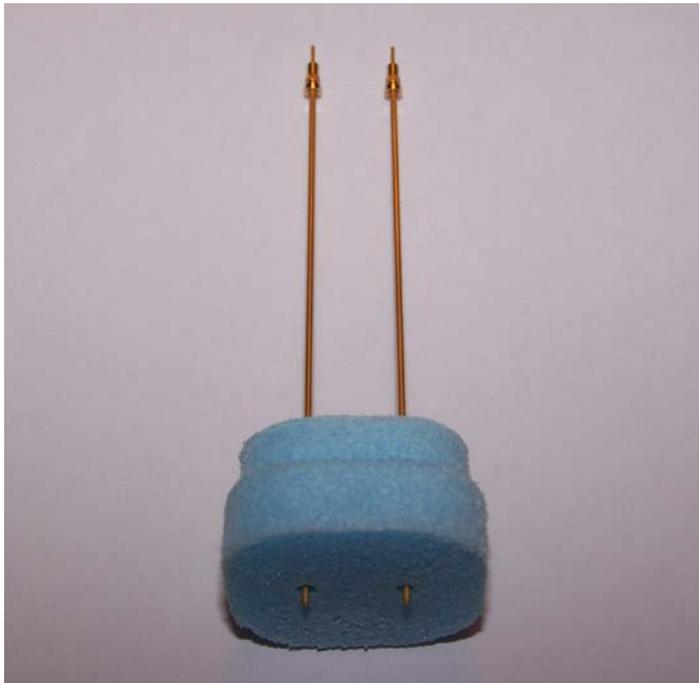
Progress – Prototype ASKAP LNA pair

- Noise
 - 40 K
- Gain
 - 28 dB
- Input Impedance
 - balanced 300 Ω
 - SE Γ_{opt} close to 150 Ω
 - thankyou ASTRON for measuring noise parameters
- Frequency Range
 - 0.7 GHz to 1.8 GHz
- Transistors
 - Avago ATF-35143
- Design
 - 2 SE amplifiers + balun and equaliser



c/- Rob Shaw

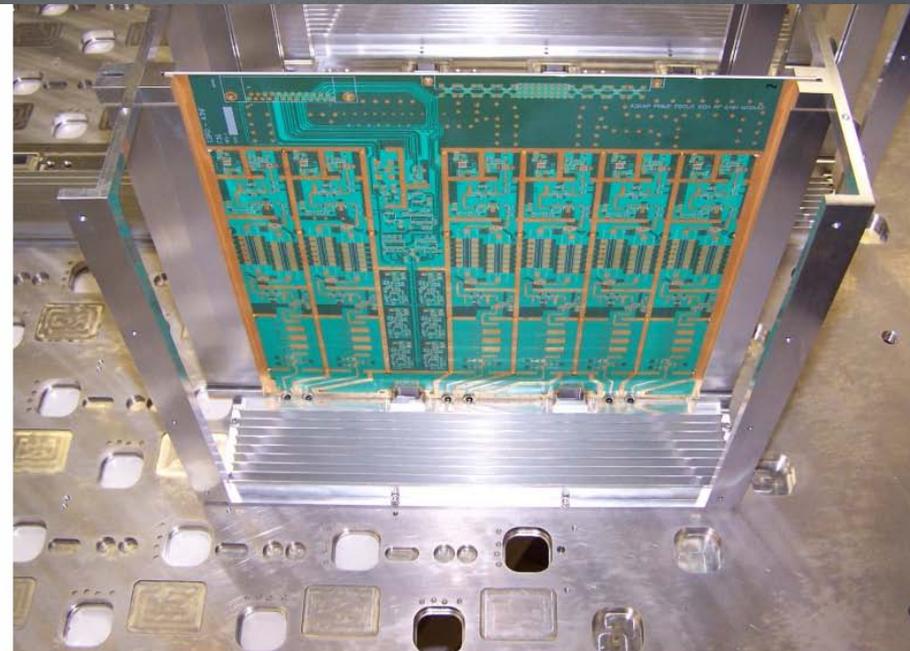
Progress – Assembly



•array feed line



•6 unit LNA shield

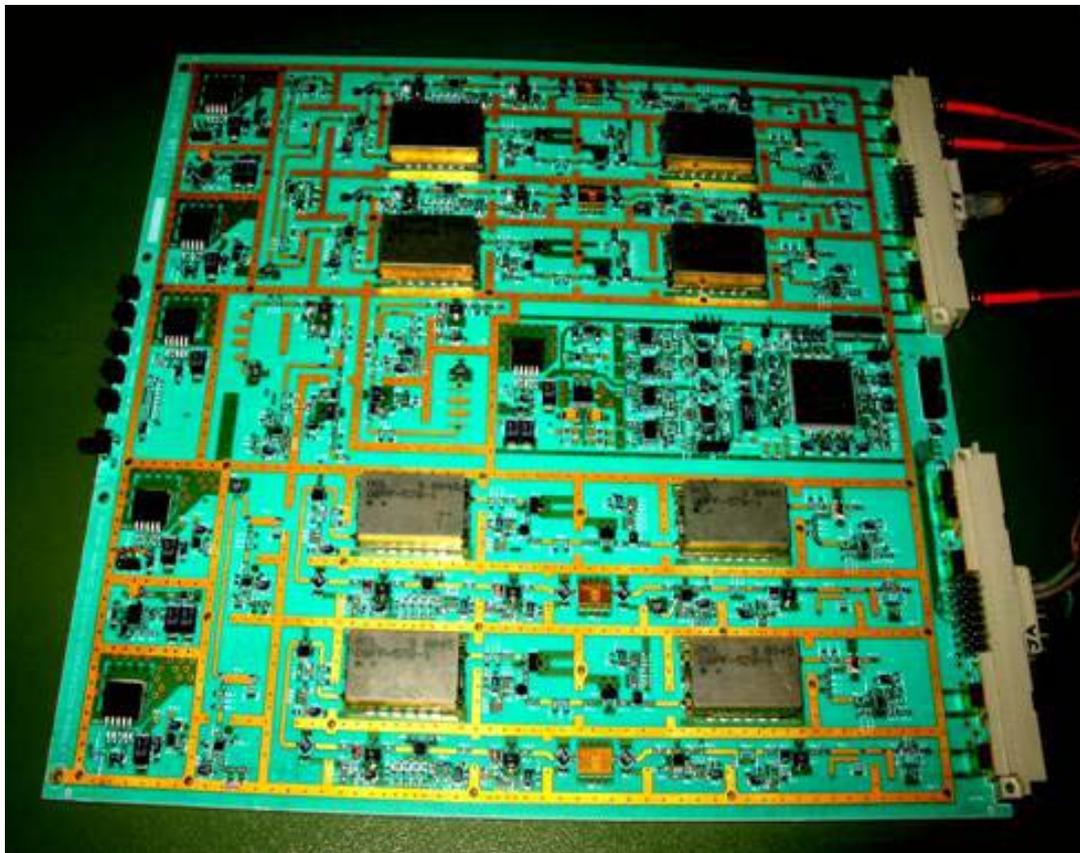


•6 unit Filter card inserts into LNA

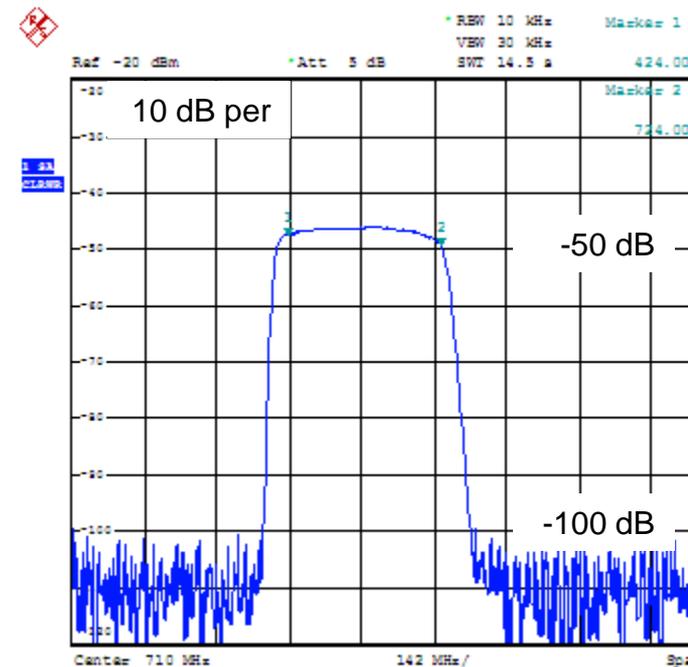


Progress – Conversion Module Prototype

- Located in shielded rack in pedestal



- passband with 40 m RF cable
- alias rejection >65 dB



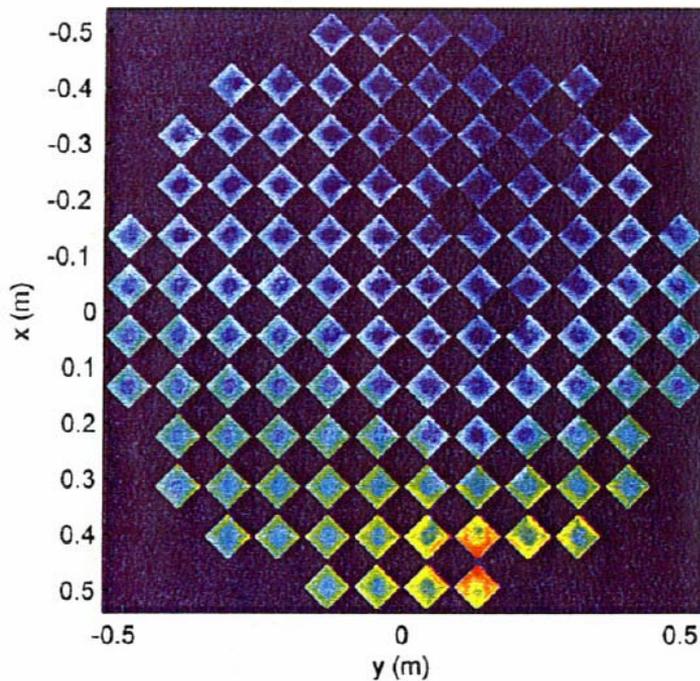
c/- K. Jeganathan

Progress – 30 deg² FOV Achieved

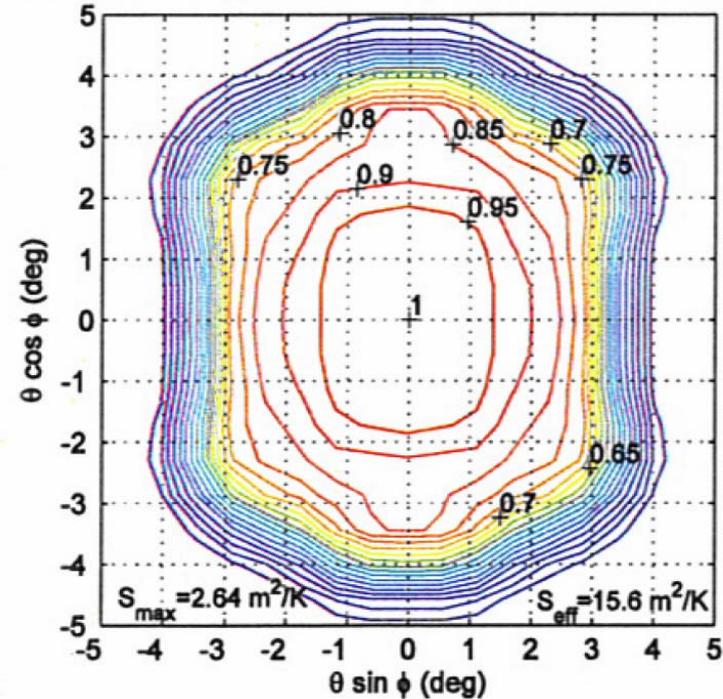
188 port ASKAP PAF



35 deg² SSFFOV



ASKAP, DSE, 1.25GHz, relative sensitivity, SSFOV=35.1deg²



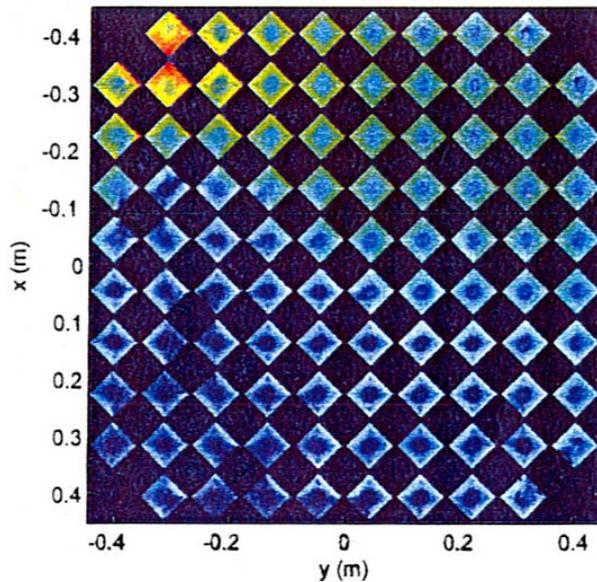
c/- Stuart Hay

Progress – 30 deg² FOV Achieved

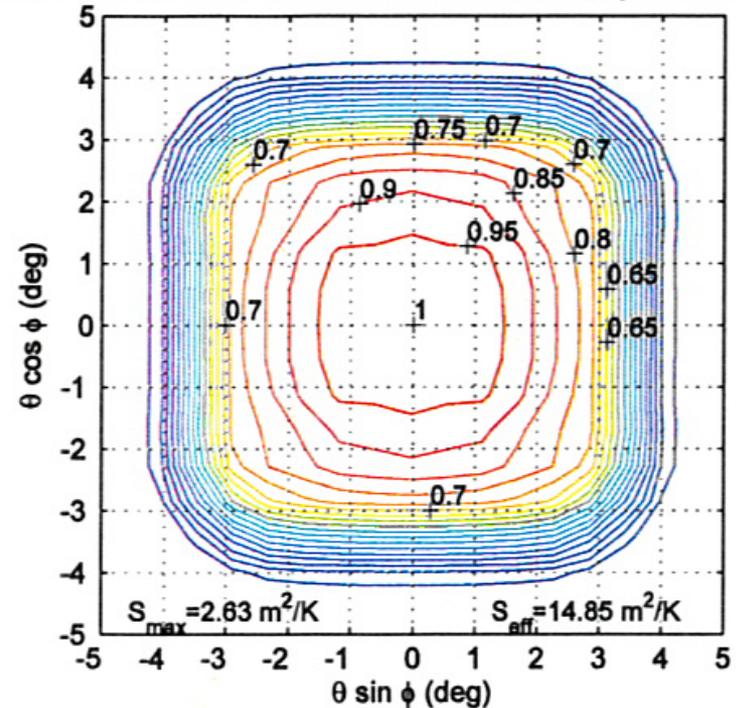
172 port ASKAP PAF



32 deg² SSFFOV



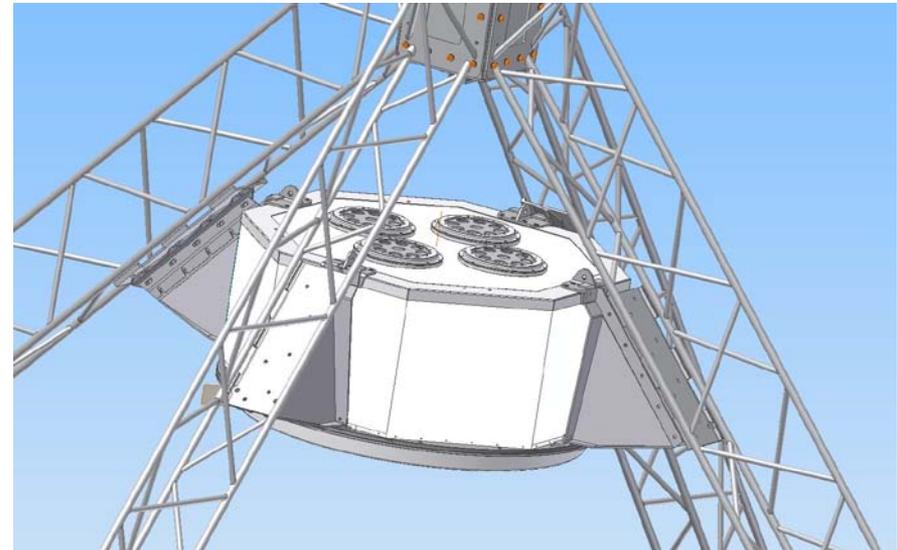
ASKAP reduced, DSE, 1.25GHz, relative sensitivity, SSFOV=32deg²



c/- Stuart Hay

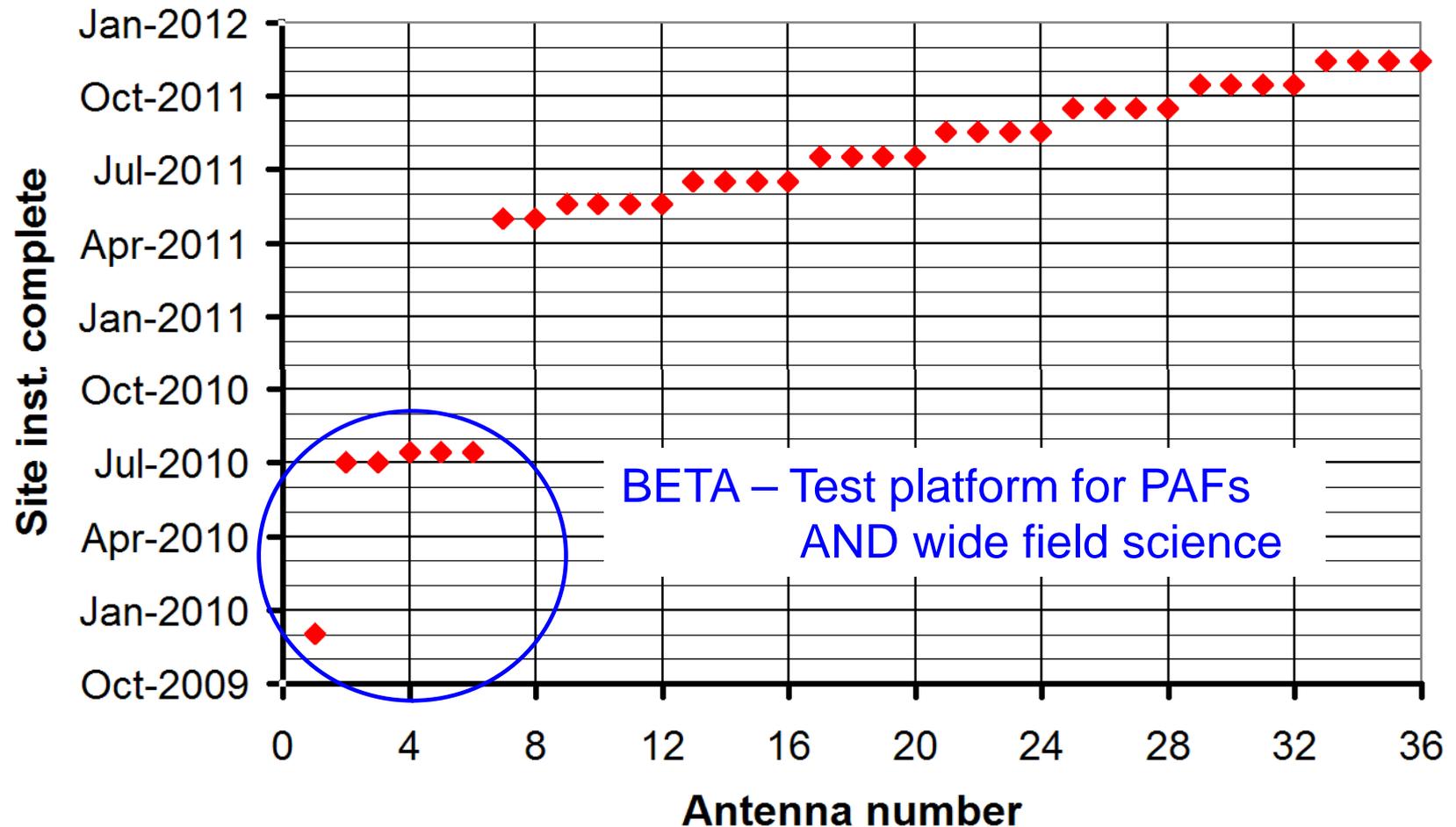
Progress – ASKAP & SKA

- Test 188 port PAF this year
 - real-time 192 port digital back-end
 - hot/cold Y-factor measurements
 - interferometry with 64 m
- PAF Port Gain Calibration
 - multi-point noise injection
 - astronomical sources
- Towards SKA
 - modelling with offset dish
 - cost reduction



Progress – Antenna Delivery Schedule

Antenna delivery



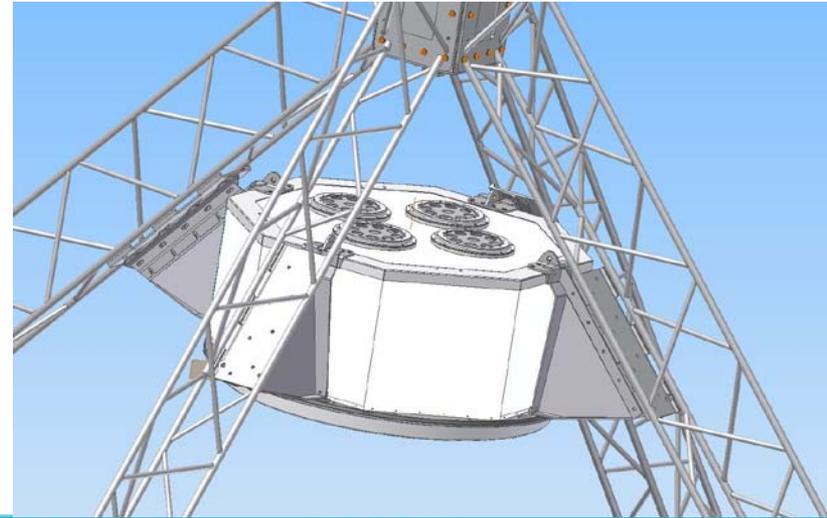
CSIRO Astronomy & Space Science

Aaron Chippendale
Engineer

Phone: +61 2 9372 4296

Email: Aaron.Chippendale@csiro.au

Web: www.atnf.csiro.au
www.csiro.au/org/CASS.html



www.csiro.au

Thank you

Contact Us

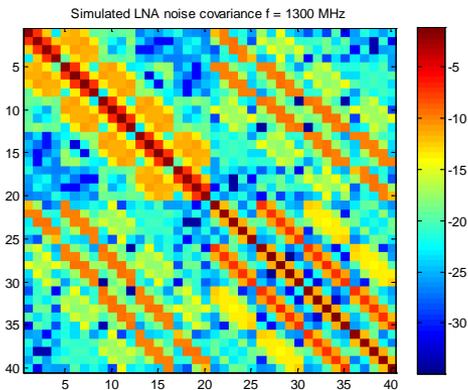
Phone: 1300 363 400 or +61 3 9545 2176

Email: enquiries@csiro.au Web: www.csiro.au

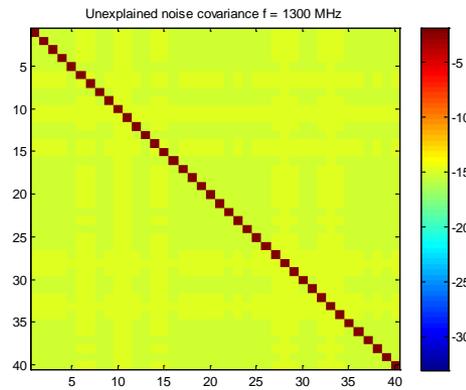


Discussion – Noise Model Components

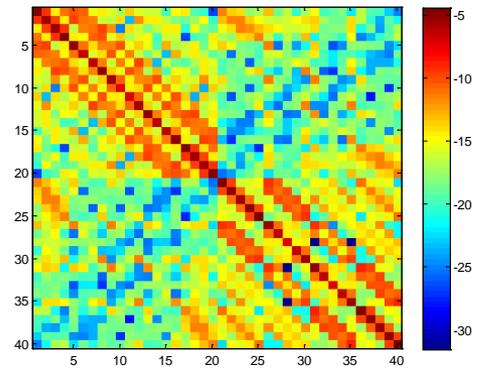
LNA noise -3dB



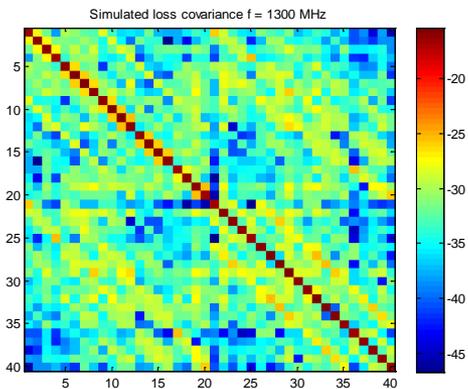
Unexplained Noise -3dB



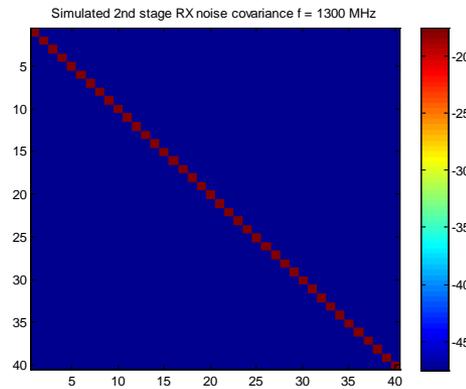
Spillover -5dB



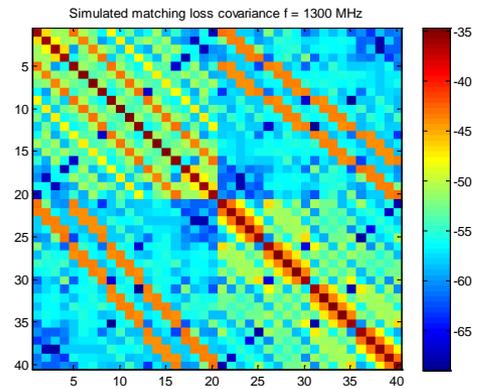
Array Losses -16dB



2nd Stage RX noise -18dB



Matching Net Loss -35dB



Discussion – Analog Connectors



Discussion – Pedestal Configuration

